

RECITATION 4

N-dimensional Gaussian distribution

Consider N Gaussian random variables $x_1, x_2, \dots, x_N$ Then joint distribution is given by the following probability density:

$$p(\vec{x}) = \frac{1}{\sqrt{(2\pi)^N \det C}} e^{-\frac{1}{2} \sum_{i,j} [(x_i - \mu_i)(x_j - \mu_j) C_{ij}^{-1}]}$$

$$\vec{x}^T C \vec{x} > 0 \quad \forall \vec{x} \neq 0$$

where μ_i is the mean of variable x_i , C is the $N \times N$ and symmetric covariance matrix such that:

positive-definite

$$C_{ii} > \sum_{j \neq i} |C_{ij}|$$

$$C_{ii} > 0$$

$$C_{ii} = \sigma_i^2 \quad \text{diagonal elements}$$

$$C_{ij} = \langle (x_i - \mu_i)(x_j - \mu_j) \rangle = \langle x_i x_j \rangle_c \quad \text{off-diagonal elements}$$

Characteristic function

$$\varphi_{\vec{x}}(\vec{k}) = \langle e^{-i \sum_i k_i x_i} \rangle = \frac{1}{\sqrt{(2\pi)^N \det C}} \int d^N \vec{x} e^{-i \vec{k} \cdot \vec{x} - \frac{1}{2} (\vec{x} - \vec{\mu})^T C^{-1} (\vec{x} - \vec{\mu})}$$

$$= \frac{1}{\sqrt{(2\pi)^N \det C}} \int d^N \vec{y} e^{-i \vec{k} \cdot \vec{y} - i \vec{k} \cdot \vec{\mu}} e^{-\frac{1}{2} \vec{y}^T C^{-1} \vec{y}} = \frac{e^{-i \vec{k} \cdot \vec{\mu}}}{\sqrt{(2\pi)^N \det C}} \int d^N \vec{y} e^{-i \vec{k} \cdot \vec{y} - \frac{1}{2} \vec{y}^T C^{-1} \vec{y}}$$

$$= \frac{e^{-i \vec{k} \cdot \vec{\mu}}}{\sqrt{(2\pi)^N \det C}} e^{-\frac{1}{2} \vec{k}^T C \vec{k}}$$

standard gaussian integral

$$= \underline{\underline{e^{-i \vec{k} \cdot \vec{\mu} - \frac{1}{2} \vec{k}^T C \vec{k}}}}$$

$$\ln \varphi_{\vec{x}}(\vec{k}) = -i \vec{k} \cdot \vec{\mu} - \frac{1}{2} \vec{k}^T C \vec{k}$$

$$\Rightarrow \begin{cases} \langle x_i \rangle_c = \mu_i \\ \langle x_i x_j \rangle_c = C_{ij} \end{cases}$$

All higher-order cumulants vanish like in the 1-D case.

Wick's theorem: if the mean of all variables vanish, i.e.

$\mu_i = 0 \quad \forall i$, then the computation of the moments simplify.

All odd moments vanish, all even moments can be written as a sum over all possible pairings of variables

$$\begin{aligned} \text{ex. } \langle x_i x_j x_k x_m \rangle &= \langle x_i x_j \rangle \langle x_k x_m \rangle + \langle x_i x_k \rangle \langle x_j x_m \rangle + \langle x_i x_m \rangle \langle x_j x_k \rangle \\ &\downarrow \\ &= C_{ij} C_{km} + C_{ik} C_{jm} + C_{im} C_{jk} \end{aligned}$$

how many pair partitions of

down elements? In how many ways we can break it into distinct pairs

$$\frac{2m!}{(2^m m!)} \stackrel{m=2}{=} \frac{4!}{2^2 2!} = \frac{24}{4 \times 2} = \frac{24}{8} = 3$$

$2m$ elements $\rightarrow$ we want to write as the product of m pairs.
there are $m!$ way to permute m pairs, so we are overcounting by a factor $m!$.

$2m!$ total number of arrangements of $2m$ elements

$m! 2^m$ - permutation of the elements in each pair
 $\downarrow$
permutation of the pairs

- Time scale separation

how to describe relaxation to equilibrium?

$$\tau_c \ll \tau_{NFP} \ll \tau_F$$

we first made a coarse grain around this scale

↳ to make the collision term tractable

From Liouville:

$$\partial_t \hat{f}_1(\vec{q}_1, \vec{p}_1, t) + \{\hat{f}_1, H_1\} = \underbrace{\frac{\partial \hat{f}_1}{\partial t}}_{\text{coll}} \quad \begin{array}{l} \text{interactions among} \\ \text{particles} \end{array}$$

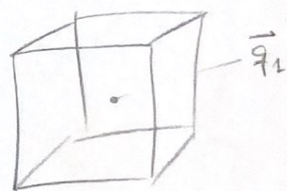
causes change $\sim \tau_{NFP}$

where $\hat{f}_1(\vec{q}_1, \vec{p}_1, t) = \frac{1}{V_c(\vec{q}_1)} \int_{V_c(\vec{q}_1)} d\vec{q}_2' \hat{f}_1(\vec{q}_2', \vec{p}_1, t)$

coarse-grained
number density

$$\{\hat{f}_1, H_1\} = \frac{\partial \hat{f}_1}{\partial \vec{q}_i} \cdot \frac{\partial H_1}{\partial \vec{p}_i} - \frac{\partial \hat{f}_1}{\partial \vec{p}_i} \cdot \frac{\partial H_1}{\partial \vec{q}_i}$$

$$H_1 = \sum_{i=1}^N \left(\frac{\vec{p}_i^2}{2m} + U(x_i) \right)$$



$\frac{\partial \hat{f}_1}{\partial t} \Big|_{\text{coll}}$ coll. changes on the time scale $\sim \tau_{NFP}$

$\hat{f}_1(t) = S\left(\frac{t}{\tau_{NFP}}\right)$ such that $S(u)$ does not depend on τ_{NFP}
 $S'(u) \sim O(\tau_{NFP}^0)$

$$\hat{f}_1'(t) = \frac{1}{\tau_{NFP}} S'\left(\frac{t}{\tau_{NFP}}\right)$$

All the dependence has been absorbed in the rescaling at t .

Using Taylor series:

$$\hat{f}_1(t+\tau) = \sum_{n=0}^{\infty} \frac{\tau^n}{n!} \frac{\partial^n \hat{f}_1}{\partial t^n} \Big|_{\tau=0} = \hat{f}_1(t) + \frac{\tau}{\tau_{NFP}} S'\left(\frac{t}{\tau_{NFP}}\right) + \frac{1}{2} \frac{\tau^2}{\tau_{NFP}^2} S''\left(\frac{t}{\tau_{NFP}}\right) + \dots$$

By choosing $\tau \ll \tau_{\text{NFP}}$ you are allowed to truncate the series.

$$\hat{A}_2(t+\tau) \approx \hat{A}_2(t) + \frac{\tau}{\tau_{\text{NFP}}} S'(\tau_{\text{NFP}})$$

$$\hat{A}_2(t+\tau) \approx \hat{A}_2(t) + \tau \hat{A}_2'(\tau)$$

$$\Rightarrow \frac{\hat{A}_2(t+\tau) - \hat{A}_2(t)}{\tau} = \hat{A}_2'(\tau)$$

$$\left. \frac{\partial \hat{A}}{\partial t} \right|_{\text{coll}} \approx \frac{1}{\tau} \left(\hat{A}_2(\vec{q}_2, \vec{p}_2, t+\tau) - \hat{A}_2(\vec{q}_2, \vec{p}_2, t) \right)$$

$$\approx \frac{1}{\tau V_c(\vec{q}_2) |d\vec{p}_2|} \underbrace{V_c(\vec{q}_2) |d\vec{p}_2| \left(\hat{A}_2(\vec{q}_2, \vec{p}_2, t+\tau) - \hat{A}_2(\vec{q}_2, \vec{p}_2, t) \right)}_{N^+(\tau, \vec{p}_2) - N^-(\tau, \vec{p}_2)}$$

$$N^-(\tau, \vec{p}_2) = \int d^3\vec{p}_2 \, d^3\theta \left(\vec{p}_2, \vec{p}_2 \rightarrow \vec{p}_2', \vec{p}_2' \right) \hat{A}_2(\vec{q}_1, \vec{p}_2, t) \hat{A}_2(\vec{q}_1, \vec{p}_2', t) |\vec{v}_2 - \vec{v}_2'| \times \tau$$

$\underbrace{\int d^3\theta \left(\vec{p}_2, \vec{p}_2 \rightarrow \vec{p}_2', \vec{p}_2' \right)}_{\text{it is an integral over } \vec{p}_2', \vec{p}_2'}$

$\times |d\vec{p}_2| \times V_c(\vec{q})$

$\#$ of collisions
 in the time interval τ which "remove"
 a particle with momentum $\vec{p}_2$

check the dimensionality:

$$p^3 \cancel{A} \frac{1}{p^3 \cancel{V}} \frac{1}{p^3 \cancel{V}} \frac{\cancel{H}}{\tau} \tau \, p^3 \cancel{V} \stackrel{ok}{=}$$

Similar computation for $N^+(\tau, \vec{p}_2)$

$$\partial_t \hat{x}_1 + \{ \hat{x}_2, H_1 \} = \int d^3 \vec{p}_2 d\theta (\vec{p}_2, \vec{p}_2 \rightarrow \vec{p}_2', \vec{p}_2') |\vec{v}_2 - \vec{v}_1| \left[\hat{x}_2(\vec{q}_1, \vec{p}_2', t) \hat{x}_1(\vec{q}_2, \vec{p}_2', t) - \hat{x}_2(\vec{q}_2, \vec{p}_2, t) \hat{x}_1(\vec{q}_1, \vec{p}_2, t) \right]$$

Rewriting:

$$\partial_t x + L_F x = C(x, x)$$

controls relaxation $\sim \tau_F$ controls relaxation $\sim \tau_{HFP}$

Using the H-theorem we have seen that on the timescale $\sim \tau_{HFP}$ when $L_F x$ does not matter, a solution to this can be expressed in terms of collisional invariants such that

$$\ln f_2 + \ln f_2 = \ln f_2' + \ln f_2'$$

5 collisional invariants:

- i) particle numbers
- ii) total momentum
- iii) energy

What about the long-term relaxation?

τ_F is the macroscopic time scale in the system

We rescale time $\hat{t} = \frac{t}{\tau_F}$ $\tau = \hat{t} \tau_F$

$$\frac{1}{\tau_F} \partial_{\hat{t}} x + \frac{1}{\tau_F} L_F x = \frac{1}{\tau_F} C(x, x) \Rightarrow \partial_{\hat{t}} x + \hat{L}_F x = \left(\frac{\tau_F}{\tau_{HFP}} \right) \hat{C}(x, x) \frac{1}{\epsilon}$$

this is why collisions stay on order $\sim O(1) + \sim O(\epsilon)$

We have adimensionalize time and we find a small parameter $\epsilon = \frac{\gamma_{HFP}}{\gamma_F}$ due to the time-scale separation.

This suggests to solve the equation perturbatively in ϵ

$$\chi = \chi_0 + \epsilon \chi_1$$

$$O\left(\frac{1}{\epsilon}\right): \hat{C}(\chi_0, t_0) = 0 \Rightarrow \chi_0 = f^{LE2}(\vec{q}, \vec{p}, t)$$

it tells that χ_0 can be parametrized by the collisional invariants! It also tells that on the timescale τ_F $\hat{C}(\chi, t_0)$ does not evolve to any longer.

$$\chi_0(\vec{q}, \vec{p}, t) = \frac{n(\vec{q}, t)}{(2\pi m k_B T)^{3/2}} e^{-\frac{(\vec{p} - m\vec{u})^2}{2m k_B T}}$$

convenient way
of parameterize it

$$\frac{n(\vec{q}, t)}{(2\pi m k_B T(\vec{q}, t))^{3/2}} e^{-\frac{m \delta \vec{v}^2}{2 k_B T}}$$

density field

$$n(\vec{q}, t) = \int d^3\vec{p} \chi_0$$

$$n \vec{u}(\vec{q}, t) = \int d^3\vec{p} \vec{v} \chi_0$$

velocity
field

$$n \frac{3}{2} k_B T(\vec{q}, t) = \int d^3\vec{p} (\vec{v} - \vec{u})^2 \frac{m}{2} \chi_0$$

temperature field

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We have derived in class the hydrodynamic equations

For these fields using the Boltzmann equation.

Let's redo together the derivation of the evolution of $\bar{u}$

$$u_\alpha = \frac{1}{n(q,t)} \int d^3\vec{p} v_\alpha f(\vec{q}, \vec{p}, t)$$

$$\partial_t u_\alpha = - \frac{1}{n^2} \partial_t n \underbrace{\int d^3\vec{p} v_\alpha f}_{u_\alpha} + \frac{1}{n(\vec{q}, t)} \int d^3\vec{p} v_\alpha \partial_t f \quad \text{using Boltzmann}$$

$$= - \partial_t n \frac{u_\alpha}{n} + \frac{1}{n} \int d^3\vec{p} v_\alpha \left[- v_\beta \cdot \frac{2}{\partial q_\beta} f \right] + \frac{1}{n} \int d^3\vec{p}_2 \int d^3\vec{p}_2' d^3\theta(\vec{p}_2', \vec{p}_2' \rightarrow \vec{p}_2, \vec{p}_1) \frac{v_\alpha^2}{v_\alpha^2}$$

$$|\vec{v}_1 - \vec{v}_2| (d\Omega' d\Omega - d\Omega d\Omega')$$

$$= - \frac{u_\alpha}{n} \partial_t n - \frac{1}{n} \partial_{q_\beta} (n \langle v_\alpha v_\beta \rangle) + \frac{1}{n} \int d^3\vec{p}_2 d^3\vec{p}_2' d^3\theta(\vec{p}_2', \vec{p}_2' \rightarrow \vec{p}_2, \vec{p}_1) |\vec{v}_1 - \vec{v}_2| (d\Omega' d\Omega - d\Omega d\Omega') v_\alpha^2$$

$$\textcircled{1} \frac{1}{2n} \int d^3\vec{p}_2 d^3\vec{p}_2' d^3\theta(\vec{p}_2', \vec{p}_2' \rightarrow \vec{p}_2, \vec{p}_1) |\vec{v}_1 - \vec{v}_2| (d\Omega' d\Omega - d\Omega d\Omega') (v_\alpha^2 + v_\alpha'^2)$$

here you are labelling $\vec{p}_2', \vec{p}_1'$ the preimage (of $\vec{p}_2, \vec{p}_1$)
but no one prevents you from calling $\vec{p}_2, \vec{p}_1$ the
preimage

$$= \frac{1}{2n} \int d^3\vec{p}_2' d^3\vec{p}_2' d^3\theta(\vec{p}_2, \vec{p}_2 \rightarrow \vec{p}_2', \vec{p}_1') |\vec{v}_1' - \vec{v}_2'| (d\Omega d\Omega' - d\Omega' d\Omega) (v_\alpha'^2 + v_\alpha^2)$$

but now, because of elastic collisions

$$d^3\vec{p}_2' d^3\vec{p}_1' = J d^3\vec{p}_2 d^3\vec{p}_1 = J d^3\vec{p}_2 d^3\vec{p}_1 = \cancel{\frac{J}{J}} d^3\vec{p}_2 d^3\vec{p}_1$$

$$d^3\theta(\vec{p}_2, \vec{p}_2 \rightarrow \vec{p}_2', \vec{p}_2') = d^3\theta(\vec{p}_2', \vec{p}_2' \rightarrow \vec{p}_2, \vec{p}_2)$$

$$|\vec{v}_1' - \vec{v}_2'| = |\vec{v}_1 - \vec{v}_2|$$

$$= \frac{1}{4n} \int d^3\vec{p}_2 d^3\vec{p}_2' d^3\theta(\vec{p}_2', \vec{p}_2' \rightarrow \vec{p}_2, \vec{p}_2) |\vec{v}_1 - \vec{v}_2| (d_2' d_2' - d_2 d_2) (v_{\alpha}^2 + v_{\alpha}^2 - v_{\alpha}^2)$$

so the collision term vanish!

Some trick using the H-theorem!

Actually, this is what we expect!

$$\text{but } \vec{p}_{cm} = \frac{\vec{p}_2 + \vec{p}_2'}{2} = \frac{\vec{p}_2' + \vec{p}_2}{2}$$

$$v_{\alpha}^2 + v_{\alpha}^2 = v_{\alpha}^2 + v_{\alpha}^2$$

Slow modes evolves because at transport, they are collisional invariants.

$$= -\frac{u_{\alpha}}{n} \partial_{\epsilon} n - \frac{1}{n} \partial_{q\beta} (n \langle v_{\alpha} v_{\beta} \rangle) + 0$$

using the first hydrodynamic equation

$$= \frac{u_{\alpha}}{n} \partial_{q\beta} (n u_{\beta}) - \frac{1}{n} \partial_{q\beta} (n \langle v_{\alpha} v_{\beta} \rangle)$$

$$\Rightarrow \underbrace{\partial_t u_{\alpha} + u_{\beta} \partial_{q\beta} u_{\alpha}}_{D_t u_{\alpha}} = \underbrace{u_{\beta} \partial_{q\beta} u_{\alpha} + u_{\alpha} \partial_{q\beta} u_{\beta} + \frac{u_{\alpha} u_{\beta}}{n} \partial_{q\beta} n}_{\partial_{q\beta} (u_{\alpha} u_{\beta})} - \frac{1}{n} \partial_{q\beta} (n \langle v_{\alpha} v_{\beta} \rangle)$$

$$D_t u_{\alpha} = \partial_{q\beta} (u_{\alpha} u_{\beta}) + \frac{u_{\alpha} u_{\beta}}{n} \partial_{q\beta} n - \frac{1}{n} \partial_{q\beta} (n \langle v_{\alpha} v_{\beta} \rangle)$$

$$\boxed{8} \quad \frac{1}{n} \partial_{q\beta} (u_{\alpha} u_{\beta} n) = -\frac{1}{n} \partial_{q\beta} (n \langle v_{\alpha} v_{\beta} \rangle - \langle v_{\alpha} \rangle \langle v_{\beta} \rangle n)$$

$$D_t u_\alpha = - \frac{1}{n} \partial_{q_\beta} (n \langle \delta v_\alpha \delta v_\beta \rangle_t)$$

where $\delta \vec{v} = \vec{v} - \langle \vec{v} \rangle_t$

$$\boxed{\mu D_t u_\alpha = - \frac{1}{n} \partial_{q_\beta} P_{\alpha\beta}}$$

with

$$P_{\alpha\beta} = \mu n \langle \delta v_\alpha \delta v_\beta \rangle_t$$

$$\boxed{D_t n(\vec{q}, t) = - n \partial_{q_\alpha} u_\alpha}$$

$$\boxed{\partial_t T + u_\alpha \partial_{q_\alpha} T = - \frac{2}{3nk_B} \partial_\alpha h_\alpha - \frac{2}{3nk_B} P_{\alpha\beta} u_\beta}$$

with heat flux along α

$$h_\alpha = \frac{n\mu}{2} \langle \delta v_\alpha \delta \vec{v} \rangle_t$$

$$\mu_{\alpha\beta} = \frac{1}{2} (\partial_\alpha u_\beta + \partial_\beta u_\alpha)$$

strain rate tensor

these are exact but not close!

We have seen in class how to compute

them to second order in ε ($t = t_0 + \varepsilon t_1$)

$$\boxed{\begin{aligned} h_\alpha^0 &= 0 \\ \dot{P}_{\alpha\beta}^0 &= n k_B T \delta_{\alpha\beta} \end{aligned}}$$

How to compute them to first order?

we go to higher order in the perturbation theory

$$O(1): \partial_t t_0 + L_F t_0 = \hat{C}(t_0, t_1) + \hat{C}(t_1, t_0)$$

(single time approximation)

$$\boxed{\partial_t t_0 + L_F t_0 = - t_1}$$

extra ingredient: single time approximation

$$t = t_0 + \varepsilon t_1 = t_0 \left(1 - \tau_{\text{HFF}} \left(\partial_t + \frac{\vec{q}_1}{H} \cdot \partial_{\vec{q}} \right) h(t_0) \right)$$

$$\Rightarrow g_1 = - \frac{\gamma_{HFP} H}{k_B T} \mu_{\alpha\beta} \left(\delta v_\alpha \delta v_\beta - \frac{\delta \bar{v}^2}{3} \delta_{\alpha\beta} \right) - \frac{\gamma_{HFP}}{T} \delta v_\alpha (\partial_{q_\alpha} T) \left(\frac{n}{2k_B T} \delta \bar{v}^2 - \frac{5}{2} \right)$$

you have used the zeroth order hydrodynamic equations here

Pressure:

$$\propto (1+g)$$

$$P_{\alpha\beta} = n \Pi \langle \delta v_\alpha \delta v_\beta \rangle_* = n H \underbrace{\frac{1}{n k_B T \delta_{\alpha\beta}}}_{\text{}} \langle \delta v_\alpha \delta v_\beta \rangle_0 + n H \frac{1}{\cancel{n k_B T}} \langle \delta v_\alpha \delta v_\beta g_1 \rangle_0$$

$$= - \frac{H^2 \gamma_{HFP}}{k_B T} \mu_{\gamma\delta} \left(\underbrace{\langle \delta v_\alpha \delta v_\beta \delta v_\gamma \delta v_\delta \rangle_0}_{(A)} - \frac{\delta_{\gamma\delta}}{3} \underbrace{\langle \delta \bar{v}^2 \delta v_\alpha \delta v_\beta \rangle_0}_{(B)} \right)$$

Using Wick's theorem

$$\frac{2m!}{2^m m!} = \frac{4!}{4 \cdot 2!} = \frac{24}{8} = 3$$

$$\frac{2m!}{2^m m!} = \frac{2}{2} = 1$$

$$(A) \quad \mu_{\gamma\delta} \langle \delta v_\alpha \delta v_\beta \delta v_\gamma \delta v_\delta \rangle_0 = \sum_{\delta, \gamma, \epsilon} \left[\mu_{\gamma\delta} \overbrace{\langle \delta v_\alpha \delta v_\beta \rangle}^{\delta_{\alpha\beta}} \overbrace{\langle \delta v_\gamma \delta v_\delta \rangle}^{\delta_{\gamma\delta}} + \overbrace{\langle \delta v_\alpha \delta v_\gamma \rangle}^{\delta_{\alpha\gamma}} \overbrace{\langle \delta v_\beta \delta v_\delta \rangle}^{\delta_{\beta\delta}} + \mu_{\gamma\delta} \overbrace{\langle \delta v_\alpha \delta v_\delta \rangle}^{\delta_{\alpha\delta}} \overbrace{\langle \delta v_\beta \delta v_\gamma \rangle}^{\delta_{\beta\gamma}} \right] = \left(\underbrace{3\mu_{\gamma\delta}}_{\text{divergence} = \partial_\delta \mu_\gamma} \delta_{\alpha\beta} + \mu_{\alpha\beta} + \mu_{\alpha\gamma} \right) \left(\frac{k_B T}{H} \right)^{\frac{1}{2}}$$

$$(B) \quad \mu_{\gamma\delta} \frac{\delta_{\gamma\delta}}{3} \langle \delta v_\alpha \delta v_\beta \delta v_\epsilon \delta v_\epsilon \rangle = \sum_{\delta, \gamma, \epsilon} \mu_{\gamma\delta} \frac{\delta_{\gamma\delta}}{3} \langle \delta v_\alpha \delta v_\beta \delta v_\epsilon \delta v_\epsilon \rangle =$$

$$= \frac{\partial_\gamma \mu_\gamma}{3} \left[\sum_\epsilon \langle \delta v_\alpha \delta v_\beta \delta v_\epsilon \delta v_\epsilon \rangle \right] = \frac{\partial_\gamma \mu_\gamma}{3} \left[\sum_\epsilon \overbrace{\langle \delta v_\alpha \delta v_\beta \rangle}^{\delta_{\alpha\beta}} \overbrace{\langle \delta v_\epsilon \delta v_\epsilon \rangle}^{\delta_{\epsilon\epsilon}} + \right]$$

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$$+ \sum_{\epsilon} \langle \overbrace{\delta v_{\alpha} \delta v_{\epsilon}}^{\delta_{\alpha\epsilon}} \rangle \langle \overbrace{\delta v_{\beta} \delta v_{\epsilon}}^{\delta_{\beta\epsilon}} \rangle + \sum_{\epsilon} \langle \overbrace{\delta v_{\alpha} \delta v_{\epsilon}}^{\delta_{\alpha\epsilon}} \rangle \langle \overbrace{\delta v_{\beta} \delta v_{\epsilon}}^{\delta_{\beta\epsilon}} \rangle \rangle$$

$$= \frac{\partial \mu}{\partial \epsilon} \left(\frac{k_B T}{n} \right)^2 n \left[\delta_{\alpha\beta} 3 + \delta_{\alpha\beta} + \delta_{\alpha\beta} \right] = \frac{\partial \mu}{\partial \epsilon} 5 \left(\frac{k_B T}{n} \right)^2 \delta_{\alpha\beta}$$

All in all:

$$P_{\alpha\beta} = nk_B T \delta_{\alpha\beta} - nk_B T \tau_{HFE} \left(\frac{\partial \mu}{\partial \epsilon} \delta_{\alpha\beta} + 2 u_{\alpha\beta} - \frac{5}{3} \frac{\partial \mu}{\partial \epsilon} \delta_{\alpha\beta} \right) \quad \underline{\underline{ok}}$$

Heat flow:

$$h_{\alpha} = \frac{n \pi}{2} \langle \delta v_{\alpha} \delta \bar{v}^2 (1 + g_{\alpha}) \rangle_0 = \frac{n \pi}{2} \langle \delta v_{\alpha} \delta \bar{v}^2 \rangle_0 + \frac{n \pi}{2} \langle \delta v_{\alpha} \delta \bar{v}^2 g_{\alpha} \rangle_0$$

$$= \frac{n \pi^2}{4 k_B T} \left(-\tau_{HFE} \right) \left(\frac{\partial g_{\beta}}{\partial T} \right) \langle \delta v_{\alpha} \delta \bar{v}^2 \delta v_{\beta} \delta \bar{v}^2 \rangle + \frac{n \pi}{4} 5 \frac{\tau_{HFE}}{T} \left(\frac{\partial g_{\beta}}{\partial T} \right) \langle \delta v_{\alpha} \delta v_{\beta} \delta \bar{v}^2 \rangle$$

(A) (B)

$$\textcircled{A} \langle \delta v_{\alpha} \delta v_{\epsilon} \delta v_{\epsilon} \delta v_{\beta} \delta v_{\beta} \delta v_{\gamma} \rangle = \sum_{\epsilon, \gamma} \delta_{\alpha\beta} \langle \delta v_{\alpha}^2 \delta v_{\epsilon}^2 \delta v_{\gamma}^2 \rangle$$

$$\frac{2m!}{2^m m!} = \frac{6!}{2^3 3!} = \frac{720}{8 \cdot 6} = 15$$

$$\frac{2m!}{2^m m!} = \frac{24}{4 \cdot 2} = 3$$

$$= \delta_{\alpha\beta} \langle \delta v_{\alpha}^6 \rangle + 2 \sum_{\epsilon \neq \alpha} \delta_{\alpha\beta} \langle \delta v_{\alpha}^4 \delta v_{\epsilon}^2 \rangle_{\epsilon=\alpha} + \sum_{\epsilon \neq \alpha} \delta_{\alpha\beta} \langle \delta v_{\alpha}^2 \delta v_{\epsilon}^4 \rangle_{\epsilon=\alpha}$$

$$+ \sum_{\epsilon \neq \alpha \neq \gamma} \delta_{\alpha\beta} \langle \delta v_{\alpha}^2 \delta v_{\epsilon}^2 \delta v_{\gamma}^2 \rangle =$$

$2 \times 2 \times 3$ (Coordinate) 2×3 (Combinatorics)

$$= \delta_{\alpha\beta} \left[15 + 12 + 6 + 2 \right] \left(\frac{k_B T}{n} \right)^3$$

$$= \delta_{\alpha\beta} 35 \left(\frac{k_B T}{n} \right)^3$$

$$(B) \quad \frac{n \pi^5}{4} \left(\frac{\gamma_{NFP}}{T} \right) (\partial_{q\beta} T(q, t)) \langle \delta v_\alpha \delta v_\beta \delta \tilde{v}^2 \rangle_0$$

$$\begin{aligned} \sum_{\delta} \langle \delta v_\alpha \delta v_\beta \delta v_\delta \delta v_\delta \rangle &= \sum_{\delta} \left[\underbrace{\langle \delta v_\alpha \delta v_\beta \rangle}_{\delta_{\alpha\beta}} \underbrace{\langle \delta v_\delta \delta v_\delta \rangle}_{\delta_{\delta\delta}} + \underbrace{\langle \delta v_\alpha \delta v_\delta \rangle}_{\delta_{\alpha\delta}} \underbrace{\langle \delta v_\beta \delta v_\delta \rangle}_{\delta_{\beta\delta}} \right. \\ &\quad \left. + \underbrace{\langle \delta v_\alpha \delta v_\delta \rangle}_{\delta_{\alpha\delta}} \underbrace{\langle \delta v_\beta \delta v_\delta \rangle}_{\delta_{\beta\delta}} \right] = [3 \delta_{\alpha\beta} + 2 \delta_{\alpha\beta}] \left(\frac{k_B T}{m} \right)^2 \end{aligned}$$

All in all:

$$\kappa_\alpha = \frac{n \pi^2}{k_B T^4} \left(- \frac{\gamma_{NFP}}{T} \right) (\partial_{q\beta} T) 35 \delta_{\alpha\beta} \left(\frac{k_B T}{m} \right)^3 + 5 \left(\frac{\gamma_{NFP}}{T} \right) (\partial_{q\beta} T) 5 \delta_{\alpha\beta} \left(\frac{k_B T}{m} \right)^2$$

$$= - \frac{n \pi^2 \gamma_{NFP}}{4 T} (\partial_{q\beta} T) \delta_{\alpha\beta} \left[\frac{\pi}{k_B T} 35 \left(\frac{k_B T}{m} \right)^3 - 25 \left(\frac{k_B T}{m} \right)^2 \right]$$

$$= - \frac{n \pi^2 \gamma_{NFP}}{4 T} (\partial_{q\alpha} T) 10 \left(\frac{k_B T}{m} \right)^2 = - \underbrace{\frac{5}{2} \frac{n \gamma_{NFP}}{m} k_B^2 T}_{\text{thermal conductivity}} (\partial_{q\alpha} T)$$